XMM-Newton CCF Release Note

XMM-CCF-REL-394

Astrometry: time variable boresight - 2023 update

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1 CCF components

| Name of CCF | VALDATE | List | of | Blocks | CAL VERSION | XSCS flag |
|--------------------|---------------------|--------------|-----|--------|-------------|-----------|
| | | change | d | | | |
| XMM_BORESIGHT_0033 | 2000-01-01T00:00:00 | OM_ANGVAR | | | | No |
| | | EMOS1_ANGVAR | | | | |
| | | EMOS2_ANGVAR | | | | |
| | | EPN_A | NGV | VAR | | |
| | | RGS1_ | ANG | VAR | | |
| | | RGS2_ | ANG | VAR | | |

2 Changes

The XMM-Newton Time Variable Boresight was implemented in 2012. It is described in the release notes XMM-CCF-REL-286 and XMM-CCF-REL-290.

The extrapolations made to derive corrections to the Euler angles based on past data imply that new updates of these corrections may be necessary from time to time. Several updates have been made since 2014, the latest being in early 2022, each update taking into account new data obtained after the previous one. They were implemented in XMM_BORESIGHT_00XX.CCF (XX=24-32), the latest being described in XMM-CCF-REL-387).

The existing CCF elements allow an extrapolation of the offset trend to be made but as new observations arrive, after some time we witness systematic deviations from the predicted offsets, particularly for the EPIC instruments. Therefore it is necessary to produce a new update using the most recent data.

| Instrument/ | P_1 | P_2 | P_3 | P_4 | P_5 | P_6 | P_7 | P_8 | P_9 | P_{10} | P ₁₁ |
|-------------|-------|---------------------------|------------------------|----------------------------|-------------------------|------------|--------|-------|-------|----------|-----------------|
| coordinate | | | | | | | | | | | |
| EPIC/Y | +0.53 | -3.1×10^{-4} | $+6.9 \times 10^{-8}$ | -3.6×10^{-12} | | +0.18 | -3.57 | 362.2 | -0.09 | -2585.85 | 4632.1 |
| EPIC/Z | +0.47 | -1.1×10^{-3} | $+2.4 \times 10^{-7}$ | -1.5×10^{-11} | | +1.37 | -11.96 | 365.4 | +0.25 | -245.85 | 3599.2 |
| OM/X | -1.48 | $+8.11 \times 10^{-4}$ | 3.94×10^{-8} | -3.16×10^{-11} | $+2.27 \times 10^{-15}$ | -1.01 | -20.91 | 365.9 | | | |
| OM/V | -2.78 | $\pm 3.10 \times 10^{-3}$ | -8.46×10^{-7} | $\pm 1.17 \times 10^{-10}$ | -5.85×10^{-15} | ± 0.77 | -17.51 | 365.0 | | | |

Table 1: Best-fit parameters implemented in this CCF.

As previously, we have analyzed the astrometry offsets derived from the pipeline PPS source lists for the EPIC and OM instruments, adding to the previous data set the observations obtained until November 2022 (Rev. 4202) for both OM and EPIC. We have modeled the offset variations with time by means of long term variations plus a periodic (nearly one year) oscillation (Talavera & Rodríguez-Pascual [1]). As in XMM-CCF-REL-387, for OM the long-term trend is characterised by a 4th order polynomial. The long-term trend in the EPIC data is, as for the last release, represented by a third order polynomial plus a periodic variation with a timescale of order 10 years. The general functional form is

$$\Delta = (P_1 + P_2 \times T + P_3 \times T^2 + P_4 T^3) + P_5 \times \cos[2\pi \times (T - P_6)/P_7)] + P_8 \times \cos[2\pi \times (T - P_9)/P_{10})]$$

where Δ is the measured offset and T is the time in Julian days elapsed since January 1, 2000 (MJD 51544.0). The new best-fit parameters are given in Table 1.

The long-term cyclic component in the EPIC data is of low amplitude (0.09 and 0.25 arcseconds in the Y and Z axes, respectively).

To avoid large deviations in the extrapolation, for EPIC we have used the IDL function TS_FCAST(X,P,N), where X are the fitted values up to November 19 2022, and P=838 and N=419.

As explained in XMM-CCF-REL-290, the same offsets obtained for EPIC can be used to process RGS data.

3 Scientific Impact of this Update

The release notes XMM-CCF-REL-286 and XMM-CCF-REL-290 explain, in detail, the improvements in the astrometry achieved with the Time Variable Boresight.

Although the corrections derived are very small, we annually update the model parameters to minimise the growth of significant deviations in the near (~ 1 year) future (but note ??).

We show in Figures 1 and 2 the offsets and the fitted corrections. The differences between XMM_BORESIGHT_0032.CCF and this new CCF can be seen there.



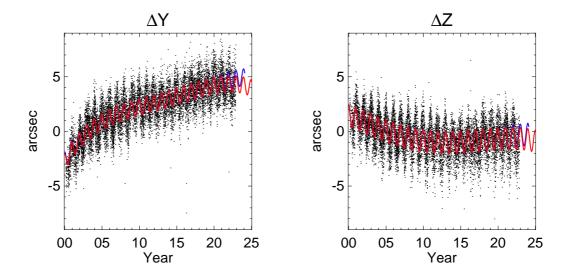


Figure 1: OM measured offsets and fit: in blue CCF_0032, in red CCF_0033.

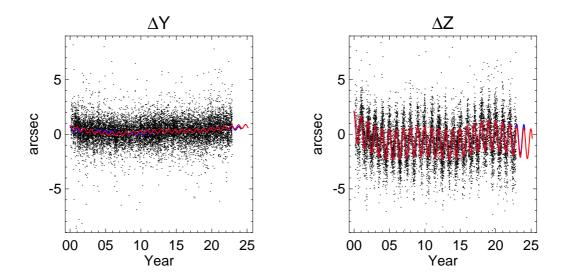


Figure 2: EPIC measured offsets and fit: in blue CCF_0032, in red CCF_0033



4 Estimated Scientific Quality

The quality of the corrections can be assessed by comparing the catalogue offsets obtained with the constant and the new variable boresight. This comparison was presented in the previous release notes, XMM-CCF-REL-286 and XMM-CCF-REL-290.

5 Test procedures

The concept of time variable boresight and its implementation were intensively tested in their first issue. At that time more than 4000 observations obtained since the beginning of the XMM-Newton operational life were processed with SAS using the new concept CCF.

Since this new release implements just a small increment in the variation of the Euler angles offsets, we have processed a recent ODFs to confirm the normal functioning of the related SAS tasks.

6 Summary of the test results

As noted before, the results of extensive tests on the time-variable boresight approach can be seen in XMM-CCF-REL-286 and XMM-CCF-REL-290. For the latest update, the tests are limited to confirming that the new boresight is correctly implemented.

7 Expected updates

The fit to the long term trend observed in the measured offsets permits an extrapolation beyond the available data. This update provides offsets until February 2025. However, following our experience these offsets will deviate from the true trend in about one year - the OM data in the Y axis (fig 1) indicate a predicted long-term trend that, by epoch 2024.0, is ~ 0.8 arcsecs lower than predicted by the previous update, driven by the apparent flatteing of the most recent OM offset measurements. Therefore we shall continue monitoring the offsets in the future to confirm the predicted trend or to modify the fit as we have been doing with the last updates.

References

[1] Talavera A., Rodríguez-Pascual P., 2011, XMM-SOC-TN-0041, available at:

http://xmm2.esac.esa.int/~xmmdoc/CoCo/CCB/DOC/Attachments/INST-TN-0041-1-0.pdf.