

XMM-Newton CCF Release Note

XMM-CCF-REL-330

Astrometry: time variable boresight. A new update.

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1 CCF components

Name of CCF	VALDATE	List of Blocks changed	CAL VERSION	XSCS flag
XMM_BORESIGHT_0025	2000-01-01T00:00:00	OM_ANGVAR EMOS1_ANGVAR EMOS2_ANGVAR EPN_ANGVAR RGS1_ANGVAR RGS2_ANGVAR		No

2 Changes

The XMM-Newton Time Variable Boresight was implemented in 2012. It is described in the release notes XMM-CCF-REL-286 and XMM-CCF-REL-290.

The extrapolations made to derive corrections to the Euler angles based on past data imply that new updates of these corrections may be necessary from time to time. An update was made in 2014 taking into account data obtained since 2012. It was implemented in XMM_BORESIGHT_0024.CCF (see XMM-CCF-REL-315).

We have observed a deviation in the offsets trend with respect to the last update predictions, particularly for the EPIC instruments. Therefore we present here a new update using the most recent data.

As we did before, we have analyzed the astrometry offsets derived from the pipeline PPS sources lists for the EPIC and OM instruments adding to the previous data set the observations obtained till August 2015. We have modeled the offsets variation with time by means of a long term variation

Table 1: Best-fit parameters implemented in this CCF.

Instrument/coordinate	P ₁	P ₂	P ₃	P ₄	P ₅	P ₆	P ₇
EPIC/Y	0.65	-5.3×10^{-4}	1.4×10^{-7}	0.18×10^{-12}	0.18	-12.37	363.7
EPIC/Z	0.65	-1.5×10^{-3}	4.5×10^{-7}	-4.3×10^{-11}	1.37	1.08	363.2
OM/X	-1.53	9.9×10^{-4}	-9.5×10^{-8}	0.0	-1.13	1.00	362.5
OM/Y	-2.39	2.1×10^{-3}	-2.0×10^{-7}	0.0	0.84	-3.10	362.6

plus a periodic (nearly one year) oscillation (Talavera & Rodríguez-Pascual [1]).

To avoid large deviations in the extrapolation beyond August 2015, for EPIC we have used the IDL function TS_FCAST(X,P,Nval), where X are the fitted values up to August 2015 (Rev. 2878), P=846 and N=423.

As explained in XMM-CCF-REL-290, the same offsets obtained for EPIC can be used to process RGS data.

$$\Delta = (P_1 + P_2 \times T + P_3 \times T^2) + P_4 T^3 + P_5 \times \cos[2\pi \times (T - P_6)/P_7]$$

where Δ is the measured offset and T is the time in Julian days elapsed since January 1, 2000. The new best-fit parameters are given in Tab.1. As before, the long term component for the EPIC is fitted with a third order polynomial, while for OM the polynomial is of second order.

3 Scientific Impact of this Update

The release notes XMM-CCF-REL-286 and XMM-CCF-REL-290 explain in detail the improvements in the astrometry achieved with the Time Variable Boresight.

We have made a small adjustment to compensate for the deviation observed, in particular for the EPIC and RGS where the corrections derived using XMM_BORESIGHT_0024.CCF differed more than 2 arc sec in the most recent observations.

We show in Figures 1 and 2 the offsets and the fitted corrections. The differences between XMM_BORESIGHT_0024.CCF and this new CCF can be seen there.

4 Estimated Scientific Quality

The quality of the corrections can be assessed by comparing the catalog offsets obtained with the constant and the new variable boresight. This comparison was presented in the previous release notes, XMM-CCF-REL-286 and XMM-CCF-REL-290.

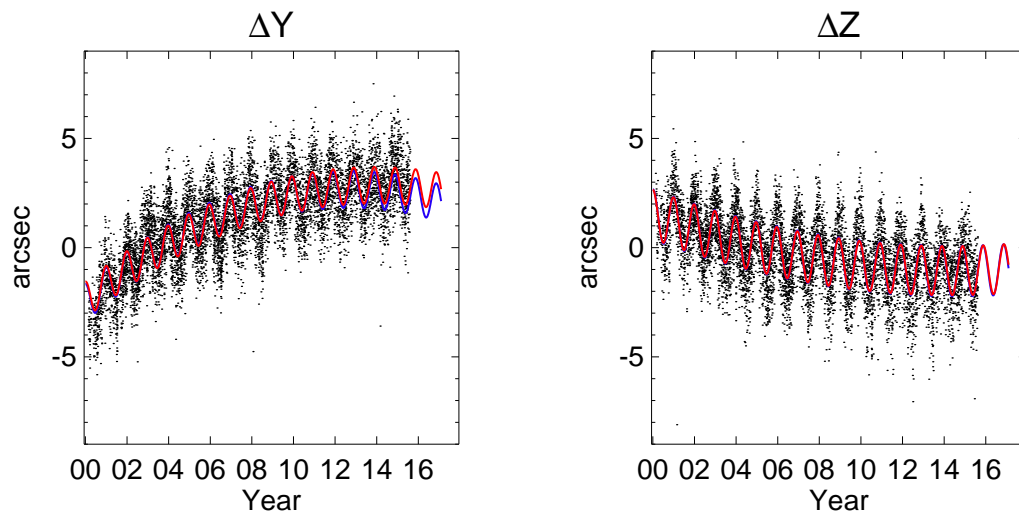


Figure 1: OM measured offsets and fit: in blue CCF_0024, in red CCF_0025

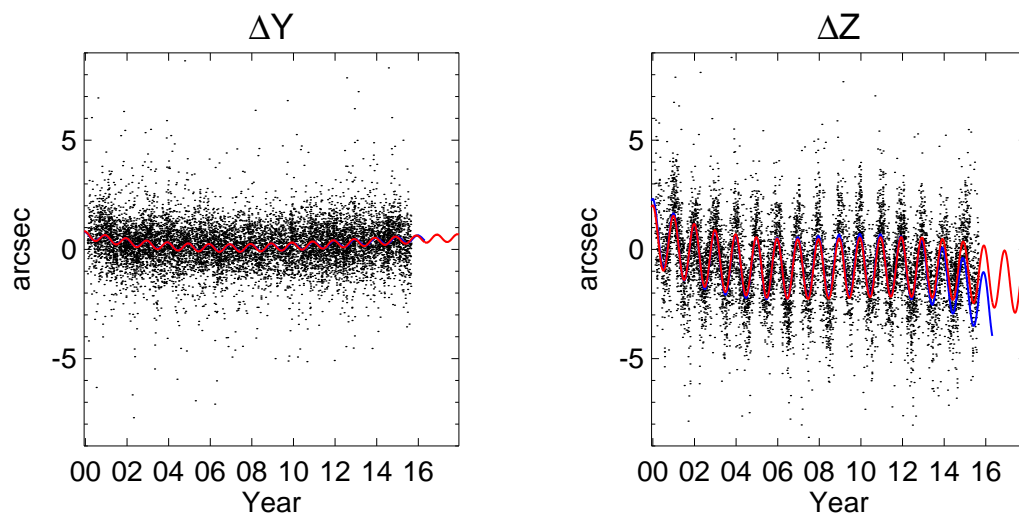


Figure 2: EPIC measured offsets and fit: in blue CCF_0024, in red CCF_0025

5 Test procedures

The concept of time variable boresight and its implementation were intensively tested in their first issue. At that time more than 4000 observations obtained since the beginning of the XMM-Newton operational life were processed with SAS using the new concept CCF.

Since this new release implements just a small increment in the variation of the Euler angles offsets, we have processed only a couple of recent ODFs to confirm the normal functioning of the related SAS tasks.

6 Summary of the test results

As said before, the tests results can be seen in XMM-CCF-REL-286 and XMM-CCF-REL-290.

7 Expected updates

The fit to the long term trend observed in the measured offsets assumes an extrapolation beyond the available data. Therefore we shall continue monitoring the offsets in the future to confirm the predicted trend or to modify the fit as we have done now.

References

[1] Talavera A., Rodríguez-Pascual P., 2011, XMM-SOC-TN-0041, available at:

<http://xmm2.esac.esa.int/~xmmdoc/CoCo/CCB/DOC/Attachments/INST-TN-0041-1-0.pdf>.