#### XMM-Newton CCF Release Note

#### XMM-CCF-REL-113

# **EPIC Spectral Response Distribution**

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### 1 CCF components

Name of CCF	VALDATE	List of Blocks changed	CAL VERSION	XSCS flag
EMOS1_REDIST_0012.CCF	2000-01-01	CCD_REDISTRIBUTION-n		NO
EMOS2_REDIST_0012.CCF	2000-01-01	CCD_REDISTRIBUTION-n		NO
EPN_REDIST_0007.CCF	2000-01-01	EBINS		NO

### 2 Changes

The MOS files have been updated with the redistribution functions from LUX. These are  $rmf\_110\_v7.2.dat$  for MOS-1 and  $rmf\_110\_v7.3.dat$  for MOS-2. The predominant changes are to the shape of the low-energy redistribution function for single-pixel events and to the low-energy tail for bi-pixel events.

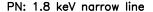
The number of energies in the PN file has been increased to 1319. This is reflected in the current batch of canned PN matrices produced by MPE.

# 3 Scientific Impact of this Update

The changes allow the MOS low-energy response to be modelled more accurately than before. The increased number of energy bins allow line features to be modelled more accurately in PN spectra.

## 4 Estimated Scientific Quality

See [1] and [2].



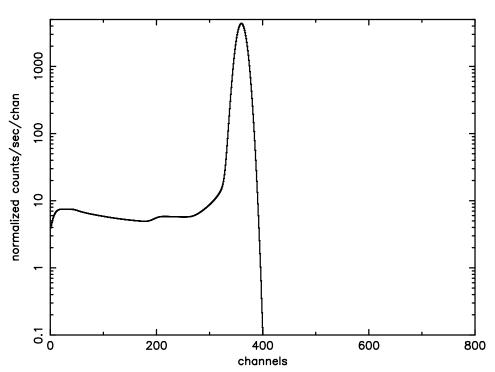


Figure 1: Comparison of PN canned response function v SAS equivalent for a 1.8 keV narrow line.

### 5 Expected Updates

The MOS redistribution files are likely to change again as the calibration improves.

## 6 Test procedures

The changes introduced here are only used within rmfgen. They are designed to allow the SAS to produce RMFs which are equivalent to the LUX and MPE canned matrices. Tests will check that they are as similar as possible.

### 7 Test results

The difference between the MOS and PN standard responses and the SAS responses generated using these new CCF elements are shown in Figures 1–3. The PN and MOS-1 response are duplicated perfectly. The SAS response for MOS-2 shows a small excess in the blue wing of the curve which is probably due to different thresholding criteria.

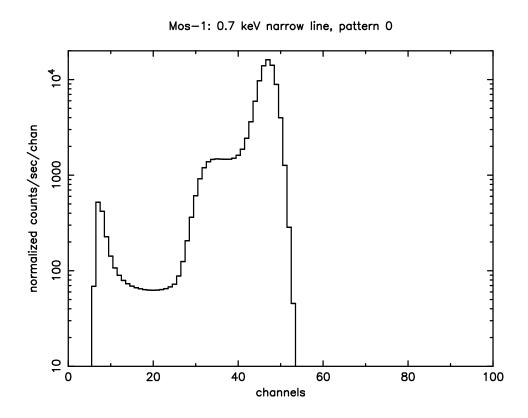


Figure 2: Comparison of MOS-1 canned response function v SAS equivalent for a  $0.7~\mathrm{keV}$  line using pattern 0.

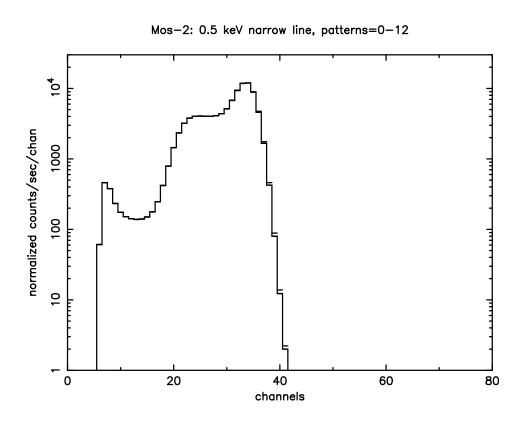


Figure 3: Comparison of MOS-2 canned response function v SAS equivalent for a  $0.5~\mathrm{keV}$  line using patterns 0--12

### References

- [1] Sembay S. et al., "Performance and Calibration of the EPIC-MOS Cameras on XMM-Newton", Proceedings of the symposium 'New visions of the X-ray Universe in the XMM-Newton and Chandra era', 26-30 November 2001, ESTEC, The Netherlands. in press.
- [2] Haberl, F., Briel, U.G., Dennerl, K. and Zavlin, V.E., "Spectral response of the EIC-PN detector: Basic dependencies", Proceedings of the symposium 'New visions of the X-ray Universe in the XMM-Newton and Chandra era'. 26-30 November 2001, ESTEC, The Netherlands. in press.